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## NONUNIQUENESS OF STEADY STATES IN ANNULAR DOMAINS FOR STREATER EQUATIONS

*Abstract.* Models introduced by R. F. Streater describe the evolution of the density and temperature of a cloud of self-gravitating particles. We study nonuniqueness of steady states in annular domains in  $\mathbb{R}^d$ ,  $d \geq 2$ .

**1. History of the problem and physical motivation.** The system we will deal with is of the form

$$\begin{aligned} (1) \quad & u_t = -\nabla \cdot \mathbf{j}, \\ & (u\theta)_t = \nabla \cdot (\lambda \nabla \theta) - \nabla \cdot (\theta \mathbf{j}) - \nabla(\phi + \phi_0) \cdot \mathbf{j}, \\ & \Delta \phi = u, \end{aligned}$$

where  $\mathbf{j} = -\kappa(\nabla u + \frac{u}{\theta}(\nabla \phi + \nabla \phi_0))$ . We will supplement this system with the following boundary conditions for  $u$  and  $\theta$ :

$$(2) \quad \frac{\partial u}{\partial \nu} + \frac{u}{\theta}(\partial_\nu \phi + \partial_\nu \phi_0) = 0,$$

$$(3) \quad \frac{\partial \theta}{\partial \nu} = 0,$$

and we will consider two kinds of boundary conditions for  $\phi$ :

$$(4) \quad \text{either the Dirichlet condition: } \phi = 0 \text{ on } \partial\Omega,$$

$$(5) \quad \text{or the “free” condition: } \phi = E_d * u_\Omega,$$

where  $u_\Omega(x) = u(x)$  in  $\Omega$  and vanishes outside  $\Omega$ , and  $(f * g)(x)$  means the convolution  $\int_\Omega f(x-y)g(y) dy$  ( $E_d$  denotes the fundamental solution of the Laplacian in  $\mathbb{R}^d$ ). We call (1) with (2)–(4) the Dirichlet problem, and (2), (3), (5) the “free” problem.

These systems describe the evolution of the density  $u(x, t) \geq 0$  and the temperature  $\theta(x, t) > 0$  in a cloud of self-gravitating particles under the

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external potential  $\phi_0(x)$  with self-induced potential  $\phi$ . The coefficients  $\kappa$  and  $\lambda$  are nonnegative functions of  $x, u, \theta$  and  $\phi$ . Furthermore, they can vanish only when  $\theta = 0$ .

In [9] this system has been derived by Streater from the Smoluchowski equation (see [8]). This was generalized in [2] to the case of gravitational and electric self-interactions.

Some results on existence and uniqueness of solutions for (1) can be found e.g. in [3] and in the references therein.

**2. Steady states equation.** Following [1] we begin by recalling basic properties of solutions to (1)–(5).

**PROPOSITION 2.1.** *For sufficiently smooth solutions of the Dirichlet and the “free” problem we have:*

- (i) *the mass  $M = \int_{\Omega} u \, dx$  and the energy  $E = \int_{\Omega} u(\theta + \phi_0 + \frac{1}{2}\phi) \, dx$  of the system are constant in time,*
- (ii) *the entropy*

$$W = \int_{\Omega} u \log \left( \frac{u}{\theta} \right) dx$$

*does not increase in time.*

*Proof.* The first assertion follows easily from the boundary conditions, the Gauss–Green theorem and (1). For the entropy, a direct calculation shows that

$$\begin{aligned} (6) \quad \frac{dW}{dt} &= \frac{d}{dt} \int_{\Omega} u \log \frac{u}{\theta} \, dx = \int_{\Omega} u_t \log \frac{u}{\theta} \, dx + \int_{\Omega} \left( \frac{u}{\theta} \right)_t \cdot \frac{\theta}{u} \, dx \\ &= - \int_{\Omega} \frac{1}{\theta} \nabla \cdot (\lambda \nabla \theta) \, dx - \int_{\Omega} \left( \frac{\nabla u}{u} + \frac{1}{\theta} (\nabla \phi + \nabla \phi_0) \cdot \mathbf{j} \right) dx \\ &= - \int_{\Omega} \lambda \frac{|\nabla \theta|^2}{\theta^2} \, dx - \int_{\Omega} \kappa u \left| \frac{\nabla u}{u} + \frac{1}{\theta} (\nabla \phi + \nabla \phi_0) \right|^2 \, dx \leq 0, \end{aligned}$$

which completes the proof. ■

For a steady state  $(u_{\infty}, \theta_{\infty})$  (i.e. a solution independent of time), the entropy of the system is constant. From this trivial fact and from (6) we deduce that

$$\nabla u_{\infty} + \frac{u_{\infty}}{\theta_{\infty}} \nabla (\phi_{\infty} + \phi_0)$$

vanishes a.e. in  $\Omega$ , and the temperature is constant:  $\theta_{\infty} = \text{const.}$  Multiplying by the nonzero factor  $e^{(\phi_{\infty} + \phi_0)/\theta_{\infty}}$  and integrating, we obtain the equation

for steady states

$$(7) \quad u_\infty = \frac{Me^{-(\phi_\infty + \phi_0)/\theta_\infty}}{\int_\Omega e^{-(\phi_\infty + \phi_0)/\theta_\infty} dx}.$$

For simplicity of notation we write  $\phi_\infty := \phi$ ,  $\theta_\infty := \theta$ ,  $u_\infty := u$ , and we will consider only the case  $\phi_0 \equiv 0$ .

As shown in [1], if we put  $\phi = \theta\psi$  in (7), the energy relation becomes

$$E = \int_\Omega u \left( \theta + \frac{1}{2}\phi \right) dx = \theta \int_\Omega u dx + \frac{\theta^2}{2} \int_\Omega u\psi dx = \theta M + \frac{\theta^2}{2} \int_\Omega \psi \Delta\psi dx.$$

Then we can replace the problem of determination of steady states with prescribed mass  $M > 0$  and energy  $E$  by the problem

$$\left( \frac{E}{M^2} \right) m^2 = m - \frac{1}{2} \int_\Omega \psi \Delta\psi dx = \mathcal{E}(m, \psi),$$

where  $m = M/\theta$  and  $\psi$  solves the Poisson–Boltzmann–Emden equation

$$(8) \quad \begin{aligned} \Delta\psi &= m \frac{e^{-\psi}}{\int_\Omega e^{-\psi} dx} && \text{in } \Omega, \\ \psi &= 0 && \text{on } \partial\Omega \end{aligned}$$

for the Dirichlet problem, and

$$(9) \quad \psi = \frac{m}{\int_\Omega e^{-\psi} dx} E_d * e^{-\psi}$$

for the “free” problem.

**3.  $k$ -symmetric solutions for the Dirichlet problem.** In this section we will deal with the two-dimensional Dirichlet problem (8) in an annular domain. Recall that the existence of radial solutions for all values of  $m > 0$  was proved in [4]. Here we show the existence of nonradial solutions. Moreover, we show nonuniqueness of solutions and we study their exact shape.

To do this we take advantage of a reasoning from [7], concerning the existence, for  $\lambda \geq 0$ , of solutions of the problem

$$(10) \quad \begin{aligned} \Delta v + \lambda e^v &= 0 && \text{in } \Omega, \\ v &= 0 && \text{on } \partial\Omega. \end{aligned}$$

Without loss of generality, we set  $\Omega = \{x \in \mathbb{R}^2 : 0 < a < |x| < 1\}$ . Let  $T_k x$  be the rotation of point  $x$  about the origin through  $2\pi/k$ , where  $k = 1, 2, 3, \dots$ , and let

$$\begin{aligned} V_k &= \{v \in H_0^1(\Omega) : v(T_k x) = v(x) \text{ a.e. in } \Omega\}, \\ V_\infty &= \{v \in H_0^1(\Omega) : v \text{ is radial}\}. \end{aligned}$$

Then  $V_k$  and  $V_\infty$  are closed subspaces of  $H_0^1(\Omega)$ .

It is evident that for  $v \in V_k$  ( $k = \infty, 1, 2, 3, \dots$ ) and  $\lambda \in \mathbb{R}$ , if  $w$  solves

$$\begin{aligned}\Delta w + \lambda e^v &= 0 && \text{in } \Omega, \\ w &= 0 && \text{on } \partial\Omega,\end{aligned}$$

then  $w$  also belongs to  $V_k$ .

Next we introduce two functionals

$$(11) \quad \Phi(v) = \int_{\Omega} e^{v(x)} dx,$$

$$(12) \quad J(v) = \frac{1}{2} \int_{\Omega} |\nabla v(x)|^2 dx,$$

and define for every  $k \in \mathbb{N}$  a subset of  $V_k$  by

$$K_{k,\mu} = \{v \in V_k : \Phi(v) = \mu\} \quad \text{for } \mu \in \mathbb{R}.$$

Let  $j_k[\mu] = \inf_{v \in K_{k,\mu}} J(v)$  whenever  $K_{k,\mu} \neq \emptyset$ . We will say that  $v$  is  $k$ -symmetric when  $v \in V_k$  and  $v \notin V_l$  for any  $l > k$ .

Our goal is to prove

**THEOREM 3.1.** *For every  $k \in \mathbb{N}$  there exists a  $k$ -symmetric solution of the Dirichlet problem (8) for some  $m > 0$ .*

In [7] the following is proved:

**THEOREM 3.2.** *For every  $k = \infty, 1, 2, 3, \dots$  such that  $K_{k,\mu} \subset V_k$  is not empty there exists  $v_\mu \in K_{k,\mu}$  which minimizes (12) (i.e.  $j_k[\mu] = J(v_\mu)$ ). Moreover, the minimizer  $v_\mu$  solves*

$$\begin{aligned}\Delta v_\mu + \lambda_\mu e^{v_\mu} &= 0 && \text{in } \Omega, \\ v_\mu &= 0 && \text{on } \partial\Omega,\end{aligned}$$

for some parameter  $\lambda_\mu \in \mathbb{R}$ . ■

We also need two technical lemmas (for details see [7, Lemmas 3.3 and 3.4]).

**LEMMA 3.3.** *For every  $k \in \mathbb{N}$  and sufficiently large  $\mu$  we have*

$$j_k[\mu] < j_\infty[\mu].$$

**LEMMA 3.4.** *For every  $k \in \mathbb{N}$  we have*

$$j_1[\mu] \leq j_2[\mu] \leq \dots \leq j_k[\mu] \leq \dots \leq j_\infty[\mu].$$

Moreover,  $j_k[\mu] < j_\infty[\mu]$  implies

$$j_1[\mu] < j_2[\mu] < \dots < j_k[\mu].$$

Notice that for  $\mu > |\Omega|$  we have  $\lambda > 0$  and the solution  $v \in K_{k,\mu}$  of the minimization problem also solves (10). Indeed, suppose, on the contrary,

that  $\lambda < 0$ . From the maximum principle for elliptic problems we have  $v \leq 0$  in  $\Omega$ . Then

$$\mu = \int_{\Omega} e^v dx \leq \int_{\Omega} 1 dx = |\Omega|,$$

which contradicts the assumption  $\mu > |\Omega|$ .

Now we recall the main theorem from [7, Th. 3.5] with its proof.

**THEOREM 3.5.** *For all  $k \in \mathbb{N}$  there exists  $\mu_k$  such that for all  $\mu > \mu_k$  the problem (10) has a  $k$ -symmetric solution. This solution satisfies*

$$\int_{\Omega} e^v dx = \mu.$$

*Proof.* By Lemma 3.3 and Theorem 3.2 there exists  $v = v_{k,\mu} \in V_k$  such that  $J(v) = j_k[\mu] < j_{\infty}[\mu]$  for some  $\mu$  sufficiently large. It suffices to show that  $v$  is  $k$ -symmetric.

Suppose, contrary to our claim, that there exists a natural number  $l > k$  such that  $v \in V_l$ . Then

$$j_l[\mu] \leq J(v) = j_k[\mu] < j_{\infty}[\mu],$$

which contradicts Lemma 3.4. ■

Now let us put  $v \mapsto -v$  in (10). Then we obtain

$$\begin{aligned} \Delta v &= \lambda e^{-v} && \text{in } \Omega, \\ v &= 0 && \text{on } \partial\Omega, \end{aligned}$$

$\Phi(-v) = \Phi_1(v) = \int_{\Omega} e^{-v} dx$ , and (8) takes the form

$$\begin{aligned} \Delta v &= \frac{m}{\Phi_1(v)} e^{-v} && \text{in } \Omega, \\ v &= 0 && \text{on } \partial\Omega. \end{aligned}$$

*Proof of Theorem 3.1.* From Theorem 3.5 we have a  $k$ -symmetric solution to the problem (10) for sufficiently large values of  $\mu = \Phi_1(v)$ . This solution solves (8) with  $m = \lambda\mu$ . ■

Since (8) has radial solutions for all  $m > 0$  (see [4]), the following corollary is obvious:

**COROLLARY 3.6.** *There is nonuniqueness of solutions for the problem (8) in a two-dimensional annulus for some values of  $m > 0$ . ■*

**4. Radial solutions for the “free” problem.** Now we will analyze (9) using dynamical systems methods. An example of use of this method is in [3].

Existence results for solutions to the problem (9) can be found in [3] and the references therein.

Our main goal in this section is to prove

**THEOREM 4.1.** *For any  $a$  and some values of  $A$  and  $m$ , the “free” problem (9) has at least two radial solutions in  $\Omega_a^A = \{x \in \mathbb{R}^d : 0 < a < |x| < A < \infty\}$  for  $3 \leq d \leq 9$ .*

Assume that  $\psi$  is radial. Abusing the notation, we set  $\psi(x) = \psi(|x|) = \psi(r)$ . Then  $\psi$  solves

$$(13) \quad \psi = \frac{m}{\int_{\Omega} e^{-\psi} dx} E_d * e^{-\psi}$$

in the annulus  $\Omega_a^A$ .

It is obvious that  $\psi$  also solves

$$\Delta\psi = m \frac{e^{-\psi}}{\int_{\Omega} e^{-\psi} dx} \quad \text{in } \Omega.$$

We set  $Q(r) = \int_{\Omega_a^r} \Delta\psi(x) dx$ . Since  $\psi$  is radial,

$$\begin{aligned} Q(r) &= \int_{\Omega_a^r} \Delta\psi(x) dx = \int_{\Omega_a^r} \left( \psi''(|x|) - \frac{d-1}{|x|} \psi'(|x|) \right) dx \\ &= \int_a^r \int_{S(s)} \left( \psi''(s) + \frac{d-1}{s} \psi'(s) \right) dS ds \\ &= \sigma_d r^{d-1} \psi'(r) - \sigma_d a^{d-1} \psi'(a). \end{aligned}$$

But  $\psi'(a) = 0$ , and furthermore, from the right hand side of (13),

$$Q(r) = \frac{m}{\int_{\Omega_a^r} e^{-\psi(x)} dx} \int_a^r \int_{S(s)} e^{-\psi(s)} dS ds = \lambda \int_a^r \sigma_d s^{d-1} e^{-\psi(s)} ds,$$

which gives

$$(14) \quad Q'(r) = \frac{d}{dr} \left( \lambda \int_a^r \sigma_d s^{d-1} e^{-\psi(s)} ds \right) = \lambda \sigma_d r^{d-1} e^{-\psi(r)}.$$

Multiplying both sides of (14) by  $r^{1-d}$  and differentiating with respect to  $r$  we obtain

$$(15) \quad (1-d)r^{-d}Q'(r) + r^{d-1}Q''(r) = -\lambda\psi'(r)\sigma_d e^{-\psi(r)}.$$

Using  $\sigma_d r^{d-1}\psi'(r) = Q(r)$  and multiplying (15) by  $r^d$ , we deduce from (14) and (15) the equation

$$(16) \quad Q''(r) + (1-d) \frac{1}{r} Q'(r) + \frac{1}{\sigma_d} r^{1-d} Q'(r) Q(r) = 0.$$

Moreover, this equation is supplemented by the boundary conditions

$$(17) \quad \begin{aligned} Q(a) &= \int \Delta \Psi(x) dx = 0, \\ Q(A) &= \int_{\Omega_a^A} \Delta \Psi(x) dx = \frac{m}{\int_{\Omega_a^A} e^{-\psi(x)} dx} \int_{\Omega_a^A} e^{-\psi(x)} dx = m. \end{aligned}$$

Notice that this procedure reduces our nonlocal problem to a local one.

Now, we can set

$$\begin{aligned} s &= \log r, \\ v(s) &= \sigma_d^{-1} r^{3-d} Q'(r), \\ w(s) &= \sigma_d^{-1} r^{2-d} Q(r). \end{aligned}$$

Then (16) will take the form of the dynamical system

$$(18) \quad \begin{cases} v'(s) = (2 - w)v, \\ w'(s) = (2 - d)w + v, \end{cases}$$

with the initial conditions

$$(19) \quad \begin{aligned} w(\log a) &= \sigma_d^{-1} a^{2-d} Q(a) = 0, \\ w(\log A) &= \sigma_d^{-1} A^{2-d} Q(A) = \sigma_d^{-1} A^{2-d} m, \end{aligned}$$

resulting from (17).

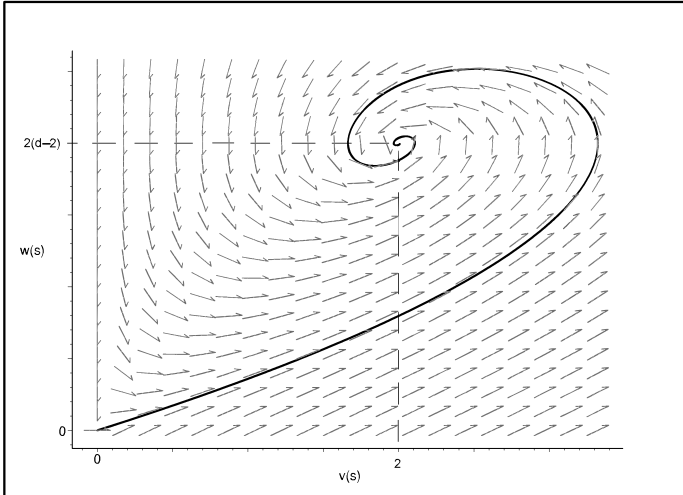


Fig. 1

Figure 1 shows the phase portrait of (18) in the first quadrant for  $d = 3$ . In the general case ( $3 \leq d \leq 9$ ) the image is similar. We have two stationary points  $P_1 = \begin{pmatrix} 0 \\ 0 \end{pmatrix}$  and  $P_2 = \begin{pmatrix} 2 \\ 2(d-2) \end{pmatrix}$ . The first one is of source type, and the second

one is a focus. These and other properties of this system can be found for example in [5].

Let us state two simple lemmas:

LEMMA 4.2. *The vector field generated by the system (18) is continuous, and transforms every interval  $\{(x, 0) : 0 < \delta < x\}$  to some curve homeomorphic to this interval. ■*

LEMMA 4.3. *For every number  $q_0 < l < q_1$  and every point  $(w_0, 0)$  there exists  $\Delta t$  such that the trajectory of the system (18), starting at time  $t_0$  from  $(w_0, 0)$ , intersects the line  $w = l$  at least  $k$  times for every fixed  $k \in \mathbb{N}$  in the time interval  $[t_0, t_0 + \Delta t]$ . ■*

Lemma 4.3 may be easily deduced from Figure 1 and the Lyapunov function for the system (18),

$$L = \frac{1}{2}(w-2)^2 + (v-2(d-2)) - 2(d-2) \log\left(\frac{v}{2(d-2)}\right).$$

*Proof of Theorem 4.1.* Since  $P_1$  is a stationary point and the vector field is continuous, for any  $t_0$ ,  $\Delta t > 0$  and  $\varepsilon_1 > 0$  we can find  $0 < \varepsilon_2 \ll 1$  such that the trajectory starting from the point  $(\varepsilon_2, 0)$  at time  $t_0$  will remain in the ball  $B((0, 0), \varepsilon_1)$  in the time interval  $[t_0, t_0 + \Delta t]$ .

For simplicity, let us fix the inner radius  $a$  of our annulus to be 1. Then from (19) we have  $t_0 = 0$ . Now, we choose a point  $(x_0, 0)$  on the  $v$ -axis and some positive number  $l_0$  such that the line  $w = l_0$  crosses the separatrix in at least two points. Next, we set the outer radius  $A$  of our annulus such that trajectory starting from  $(x_0, 0)$  for  $t_0 = 0$  intersects the line  $w = l_0$  twice in the time interval  $[0, \Delta t]$ , where  $\Delta t = \log A$ . Finally, we determine the parameter  $m_0$  such that  $\sigma_d^{-1} A^{2-d} m_0 = l_0$ .

As mentioned before we can now choose  $(\varepsilon, 0)$  such that the trajectory starting from  $(\varepsilon, 0)$  at time  $t_0$  will not cross the line  $w = l_0$  before time  $t_0 + \Delta t$ . Now by Lemma 4.2 the interval  $I = \{(x, 0) : \varepsilon < x < x_0\}$  is transformed by the dynamical system (18) to a curve  $\Gamma$ , homeomorphic to this interval. This means that there exist at least two points  $\mathcal{K}_1 = (x_1^k, l_0)$ ,  $\mathcal{K}_2 = (x_2^k, l_0)$  such that  $\Gamma$  crosses the line  $w = l_0$  at  $\mathcal{K}_1, \mathcal{K}_2$ . Hence we can find two points  $\mathcal{N}_1 = (x_1^n, 0)$ ,  $\mathcal{N}_2 = (x_2^n, 0)$  in  $I$  such that the trajectories starting from  $\mathcal{N}_1, \mathcal{N}_2$  will, after time  $t_0 + \Delta t$ , stop at line  $w = l_0$ . These are the trajectories we are looking for. They represent two different solutions of (9) on the annulus  $\Omega_a^A$  with parameter  $m = m_0$ . ■

COROLLARY 4.4. *Taking  $a, A, m$  in the proof of Theorem 4.1 more carefully, we can construct exactly  $k$  different radial solutions for the problem (9) in  $\Omega_a^A$  and any fixed  $k \in \mathbb{N}$ .*



REMARK 4.5. By applying this approach, for a fixed  $m$ , it is not difficult to show uniqueness of radial solutions when  $A/a > 0$  is sufficiently small and nonexistence of such solutions for  $A/a \gg 1$ .

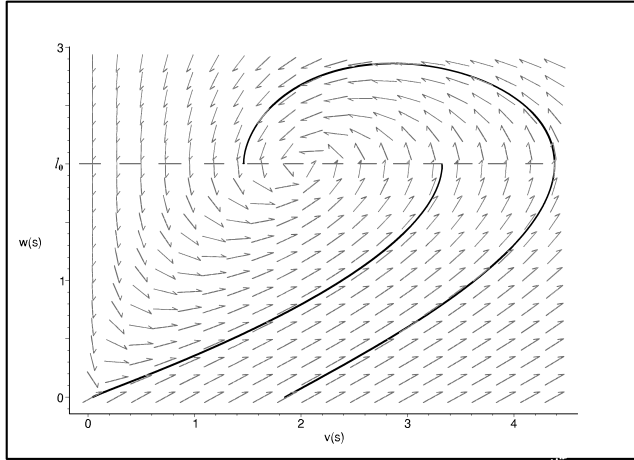


Fig. 2

Examples of such solutions can be found below. We put  $r = 1$ ,  $R = e^3$ ,  $m = 504.8424623$ . After a suitable change of variables we find two trajectories such that for  $t_0 = 0$ ,  $t_1 = 3$  we have  $w(t_0) = 0$  and  $w(t_1) = 2.00015$  (we used the Fehlberg fourth-fifth order Runge–Kutta method). In Figure 2 we have both trajectories  $(w, v)$ . The left one starts from  $(0.04026, 0)$ , and the right one from  $(1.77, 0)$ . Figure 3 shows the section of the graph of the density function along some ray starting from  $(0, 0)$ .

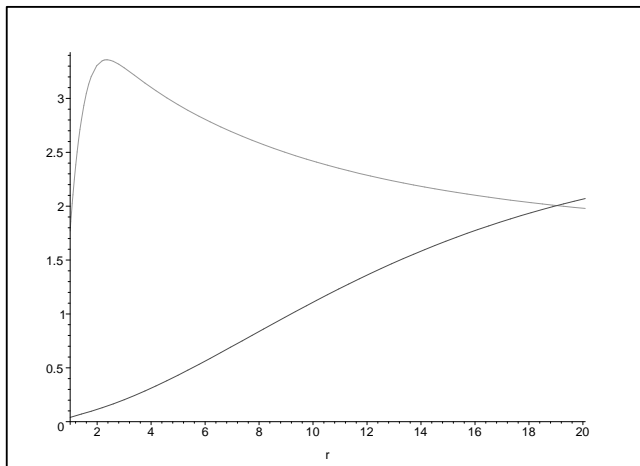


Fig. 3

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